Rotation Curve Kinematics of Galaxies in the GOODS-N Field

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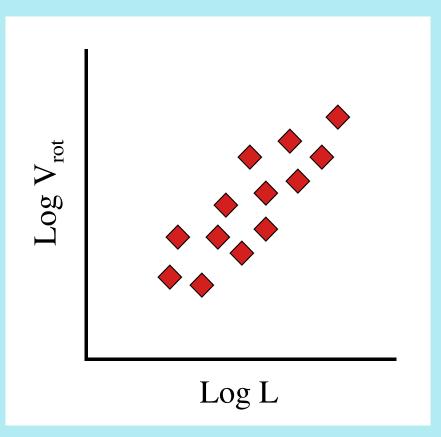
DEEP collaboration

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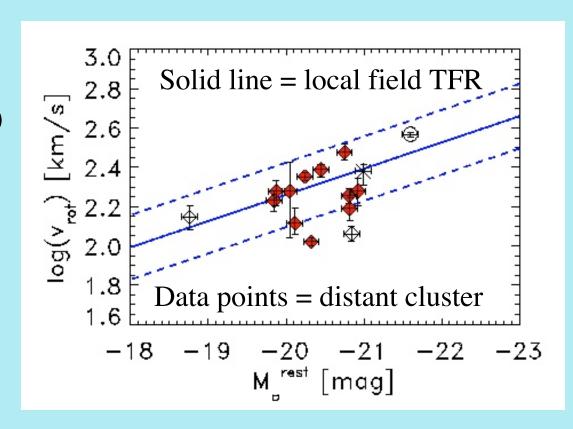
Kinematics as Evolutionary Diagnostic

- Tully-Fisher relation: comparison of disk galaxy M/L
- **Zeropoint:** luminosity evolution
- Slope: mass-segregated evolution
- Scatter: stellar populations, star formation histories



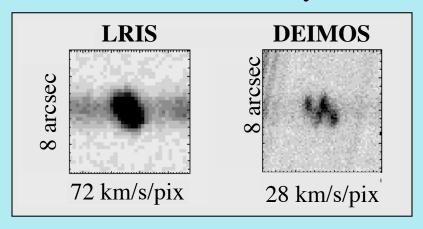
Tully-Fisher Evolution: Current Status

- Several studies show
 0.5 mag luminosity
 evolution to z=1 (e.g.
 Vogt et al. 1996, 1997)
- Others find 2 mag evolution by z=0.5 (Rix et al. 1997, Simard et al. 1998)
- Small samples and selection may be an issue (Ziegler et al. 2002)



GOODS-N Dataset

- Spectroscopy: Keck Treasury Redshift Survey
 - Keck + DEIMOS spectra of ~2000 objects with R<24.3
 - Spectra being made public at this meeting!
 - Will provide redshifts, raw velocity measurements

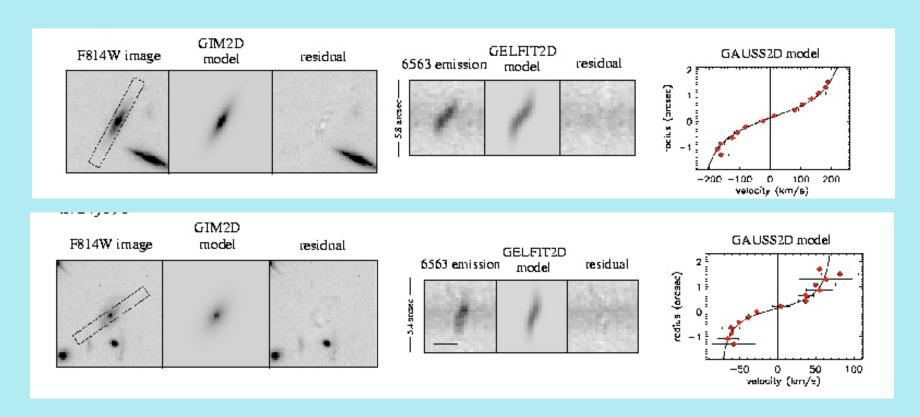


- Imaging: ACS survey images
 - Will provide restframe luminosities, colors, inclinations, position angles

1/1/04

Distant Galaxy Rotation Curve Analysis

Need to account for seeing, instrumental blurring, orientation of galaxy with respect to slit



Probing the GOODS-N Kinematic Sample

- Goal is to select galaxies in 2-3 redshift bins using specific, identical criteria (color, luminosity)
- Compare samples internally and to identicallyselected local sample



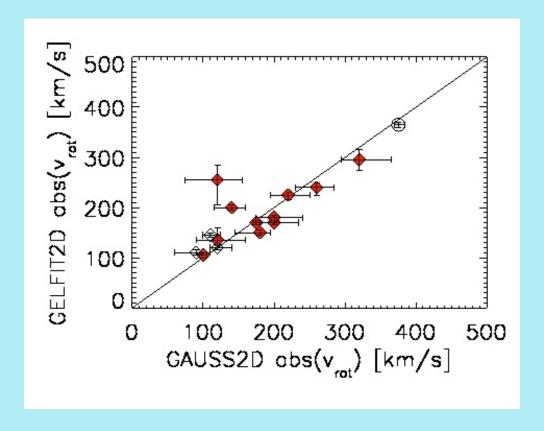


• Examined \sim 750 spectra, expect total sample of \sim 75 galaxies with z < 0.6

Further Topics to be Explored

- Extension to other density regimes: **groups**, **clusters**
- Physical basis of Tully-Fisher relation
 - Does gas fill a galaxy?
 - Do kinematic and morphological peculiarities correspond?
- Redshifting local data for comparison to distant galaxies with peculiar kinematics
- Comparing velocity measures from rotation curves and linewidths
- IFU follow-up spectroscopy: full 2D kinematics

Two Methods Compared



Very similar results!